

Keck Integral-Field Spectroscopy of M87 Reveals an Intrinsically Triaxial Galaxy and a Revised Black Hole Mass

Emily Liepold, UC Berkeley
emilyliepold@berkeley.edu

Liepold, Ma, and Walsh, ApJL, 945 L35. (12 days old!)
Download the paper at emilyliepold.com/M87

Our Observations

Our Data

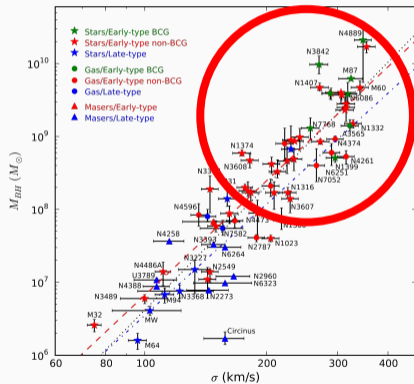
Triaxiality!

Triaxial Schwarzschild Modelling

Results!

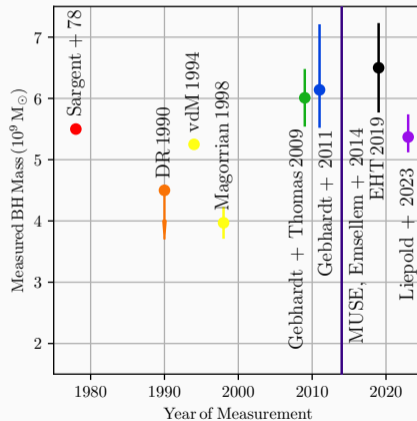
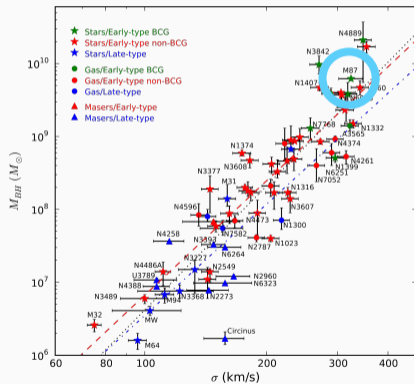
Motivation: What are we looking at?

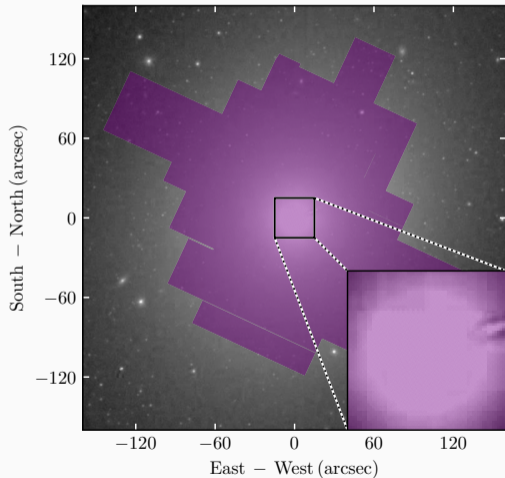
The MASSIVE Survey targets MASSIVE galaxies with MASSIVE black holes



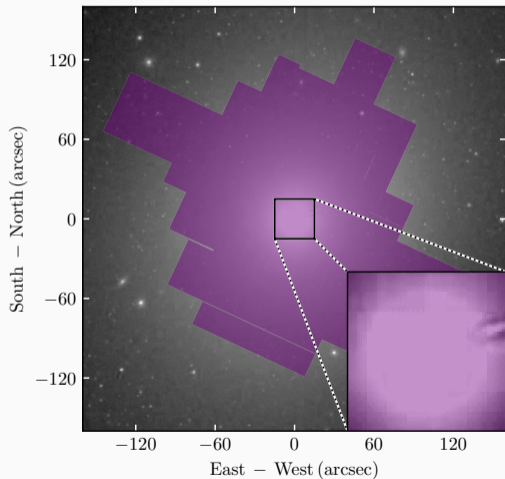
Motivation: What are we looking at?

The MASSIVE Survey targets MASSIVE galaxies with MASSIVE black holes

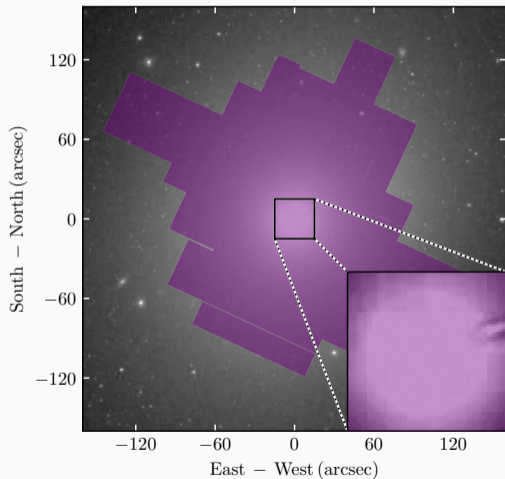




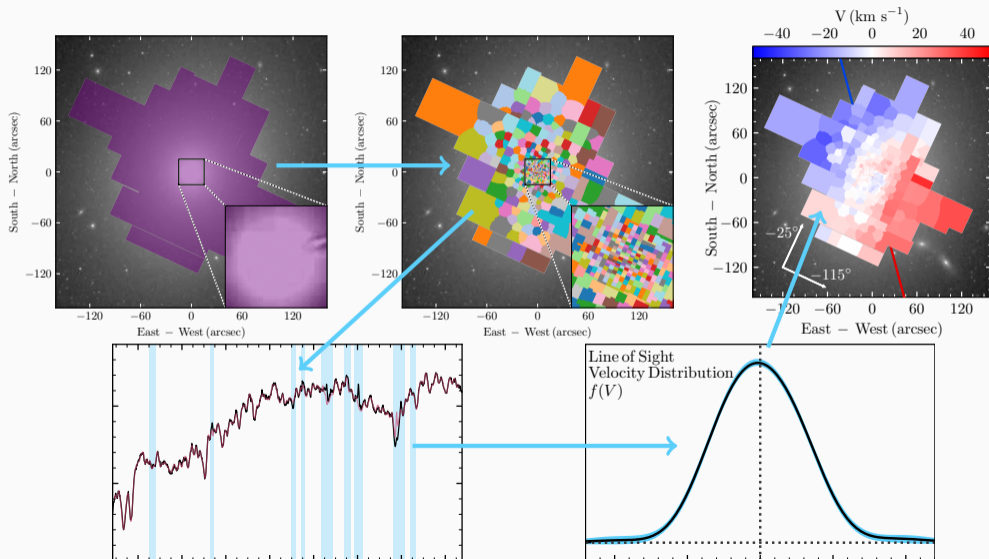
- We observed M87 with Keck Cosmic Web Imager (KCWI) during four observing runs from May 2020 - April 2022.
- This is an integral field unit, yielding a distinct spectrum at each spatial pixel.



- We observed M87 with Keck Cosmic Web Imager (KCWI) during four observing runs from May 2020 - April 2022.
- This is an integral field unit, yielding a distinct spectrum at each spatial pixel.
- 62 pointings were observed, each corresponding to a $20.4'' \times 33''$ FOV with $0.3'' \times 1.4''$ spatial pixels
- The full FOV spans about $250''$ along the photometric major axis and $300''$ along the minor (11.6 square arcmin in total!)



- We observed M87 with Keck Cosmic Web Imager (KCWI) during four observing runs from May 2020 - April 2022.
- This is an integral field unit, yielding a distinct spectrum at each spatial pixel.
- 62 pointings were observed, each corresponding to a $20.4'' \times 33''$ FOV with $0.3'' \times 1.4''$ spatial pixels
- The full FOV spans about $250''$ along the photometric major axis and $300''$ along the minor (11.6 square arcmin in total!)
- The spectra are usable from 3500\AA and 5600\AA with $R \sim 900$



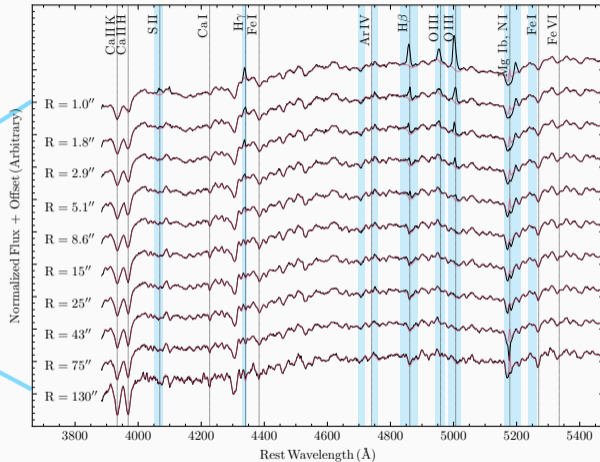
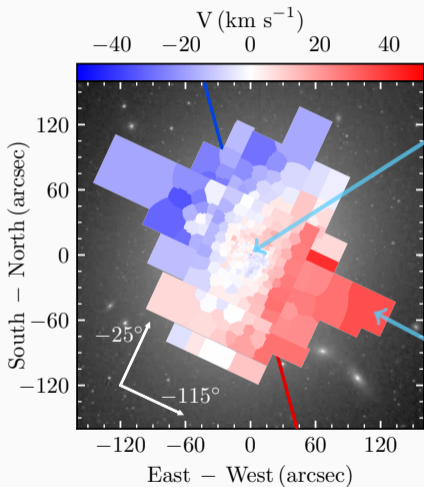
Our Observations

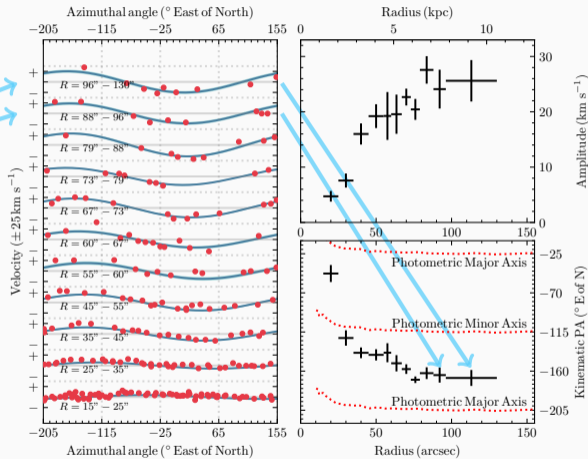
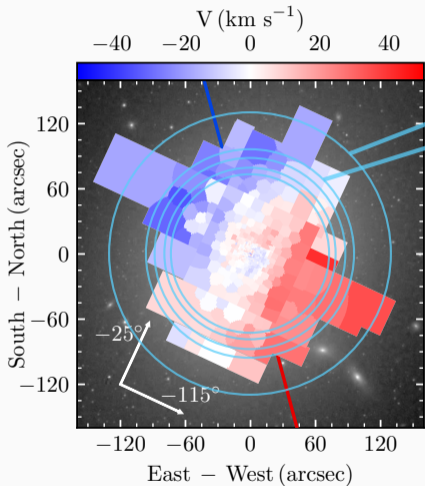
Our Data

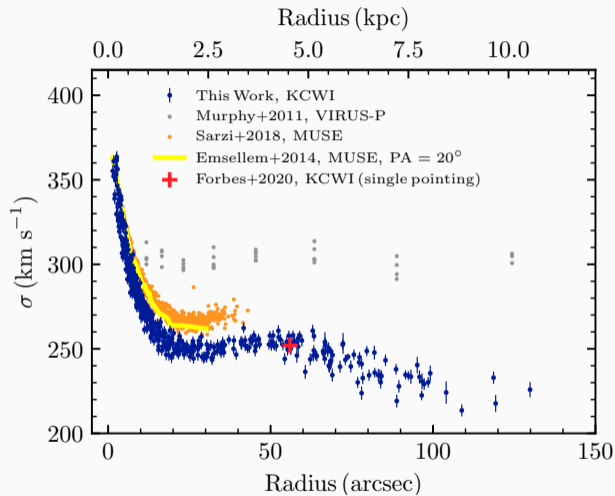
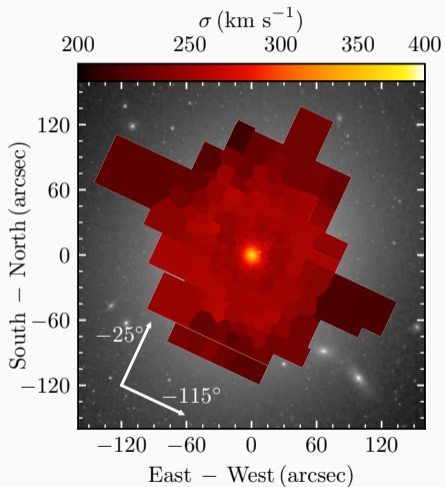
Triaxiality!

Triaxial Schwarzschild Modelling

Results!







Our Observations

Our Data

Triaxiality!

Triaxial Schwarzschild Modelling

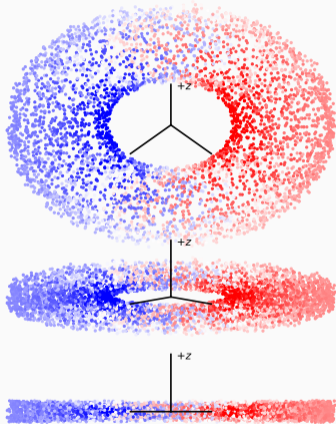
Results!

Motivation: Why do we care about the shape?

Shape of $\rho \rightarrow$ Shape of $\Phi \rightarrow$ Symmetries of $\Phi \rightarrow$ Conserved quantities and allowed orbits

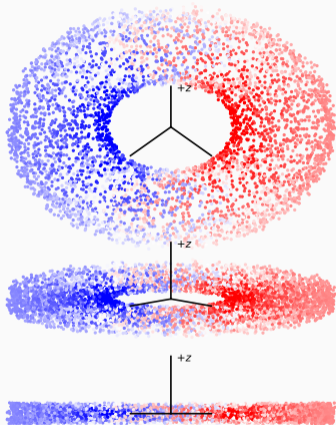
Symmetry		Conserved Quantity	Orbits
Spherical	$\frac{d\Phi}{d\Omega} = 0$	(E, \vec{L})	Rosettes in fixed planes
Axisymmetry	$\frac{d\Phi}{d\phi} = 0$	(E, L_z, I_3)	Loops about symmetry axis
Triaxiality	Eh...	(E, I_2, I_3)	It's complicated...

More Motivation: Kinematic fingerprints of non-axisymmetry



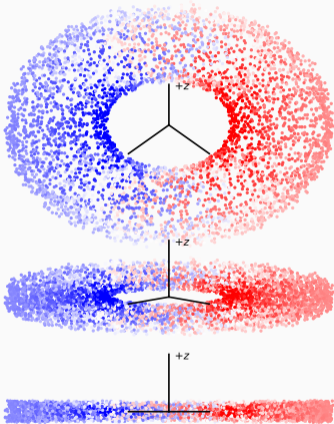
- Orbits with axisymmetric symmetry **always** have a *kinematic axis* perpendicular to the projected symmetry axis

More Motivation: Kinematic fingerprints of non-axisymmetry



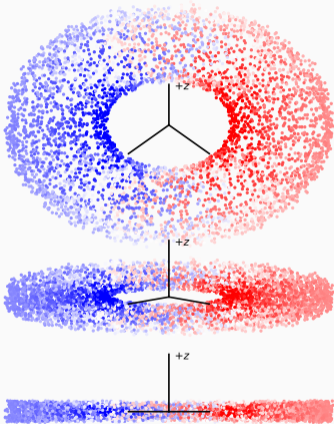
- Orbits with axisymmetric symmetry **always** have a *kinematic axis* perpendicular to the projected symmetry axis
- In oblate axisymmetry, the projected symmetry axis is perpendicular to the photometric major axis.

More Motivation: Kinematic fingerprints of non-axisymmetry



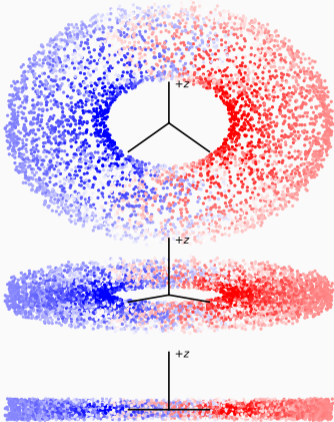
- Orbits with axisymmetric symmetry **always** have a *kinematic axis* perpendicular to the projected symmetry axis
- In oblate axisymmetry, the projected symmetry axis is perpendicular to the photometric major axis.
- Axisymmetric models **cannot** exhibit kinematic misalignment.

More Motivation: Kinematic fingerprints of non-axisymmetry



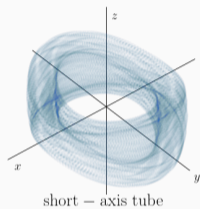
- Orbits with axisymmetric symmetry **always** have a *kinematic axis* perpendicular to the projected symmetry axis
- In oblate axisymmetry, the projected symmetry axis is perpendicular to the photometric major axis.
- Axisymmetric models **cannot** exhibit kinematic misalignment.
- Triaxial modelling (or modelling with less symmetry) is required to reproduce the velocity fields if there is kinematic misalignment or other non-bisymmetric features

More Motivation: Kinematic fingerprints of non-axisymmetry

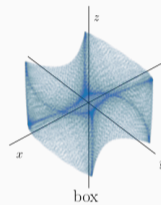


- Orbits with axisymmetric symmetry **always** have a *kinematic axis* perpendicular to the projected symmetry axis
- In oblate axisymmetry, the projected symmetry axis is perpendicular to the photometric major axis.
- Axisymmetric models **cannot** exhibit kinematic misalignment.
- Triaxial modelling (or modelling with less symmetry) is required to reproduce the velocity fields if there is kinematic misalignment or other non-bisymmetric features
- (That's M87!)

Loop Orbits

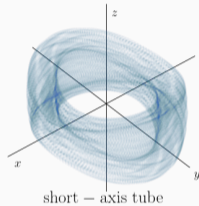


Box Orbits



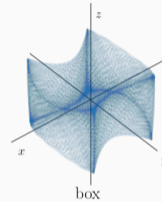
Orbits in triaxial potentials

Loop Orbits



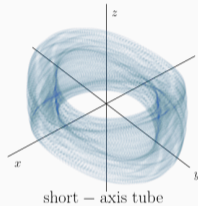
Appears in axisymmetric potentials

Box Orbits



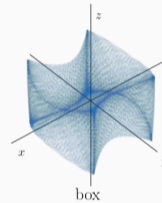
Not present in axisymmetry!

Loop Orbits



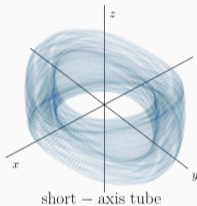
Appears in axisymmetric potentials
Persistent sense of rotation about
either the **short** or **long** axis

Box Orbits



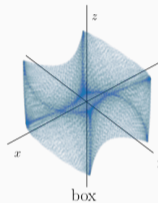
Not present in axisymmetry!
No persistent sense of rotation

Loop Orbits



Appears in axisymmetric potentials
Persistent sense of rotation about
either the **short** or **long** axis
Centrophobic

Box Orbits



Not present in axisymmetry!
No persistent sense of rotation
Can be **Centrophilic**

Our Observations

Our Data

Triaxiality!

Triaxial Schwarzschild Modelling

Results!

Schwarzschild 1979: Can triaxial stellar systems in dynamical equilibrium be self-consistent?

Strategy:

1. Propose a (triaxial) stellar density distribution
2. Integrate representative orbits that span the phase space
3. Superimpose those orbits such that (1) is reproduced

Schwarzschild 1979: Can triaxial stellar systems in dynamical equilibrium be self-consistent?

Strategy:

1. Propose a (triaxial) stellar density distribution
2. Integrate representative orbits that span the phase space
3. Superimpose those orbits such that (1) is reproduced

This turns out to be easy for reasonable proposed models. We can also try to fit kinematic observables to compare different proposed potentials.

van den Bosch+ 2008: Development of a **fortan**-based code for Schwarzschild orbit modelling in triaxial stellar potentials.

Model includes BH, stars, and dark matter halo:

$$\Phi = \Phi_{BH} + \Phi_* + \Phi_{DM}$$

Stellar kinematics (LOSVDs) described by Gauss-Hermite expansion with $y = (v - V)/\sigma$:

$$f(v) = \frac{e^{-\frac{y^2}{2}}}{\sqrt{2\pi\sigma^2}} \left[1 + \sum_{m=3}^n h_m H_m(y) \right]$$

2D (projected) and 3D (intrinsic) mass distributions are constrained for self-consistency. The code was un-named. We call our improved version 'TriOS' (**T**riaxial **O**rbit **S**uperposition)

Each **TriOS** model gives a χ^2 value for a single point in the parameter-space

- We need to search over M_{BH} , M/L (1 or 2 parameters), shape (3 parameters), and halo (1 or 2 parameters) – at least **6-8 dimensions**. (Grid Searches are inefficient)

Each **TriOS** model gives a χ^2 value for a single point in the parameter-space

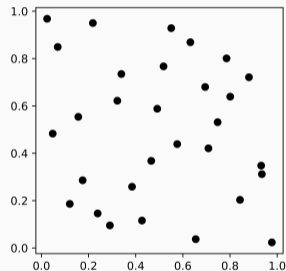
- We need to search over M_{BH} , M/L (1 or 2 parameters), shape (3 parameters), and halo (1 or 2 parameters) – at least **6-8 dimensions**. (Grid Searches are inefficient)
- This is **expensive**. Each model evaluation takes 10-30 CPU hours. (Highly iterative searches are impractical)

Each **TriOS** model gives a χ^2 value for a single point in the parameter-space

- We need to search over M_{BH} , M/L (1 or 2 parameters), shape (3 parameters), and halo (1 or 2 parameters) – at least **6-8 dimensions**. (Grid Searches are inefficient)
- This is **expensive**. Each model evaluation takes 10-30 CPU hours. (Highly iterative searches are impractical)
- As data improves, confidence volumes **shrink** with $\sim (\text{Number of Constraints})^{-D/2}$

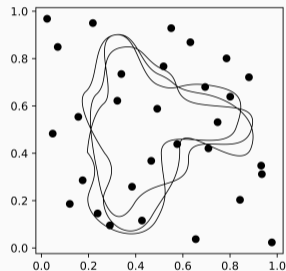
Our Strategy (inspired by Bayesian Optimization and nested sampling):

1. Sparsely populate the space



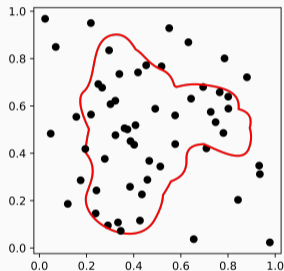
Our Strategy (inspired by Bayesian Optimization and nested sampling):

1. Sparsely populate the space
2. Use Gaussian Process regression to model the χ^2 landscape



Our Strategy (inspired by Bayesian Optimization and nested sampling):

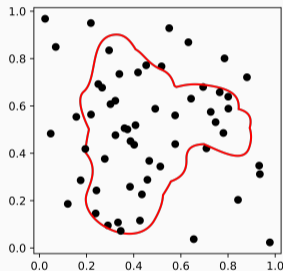
1. Sparsely populate the space
2. Use Gaussian Process regression to model the χ^2 landscape
3. Populate regions that are likely to be useful



Our Strategy (inspired by Bayesian Optimization and nested sampling):

1. Sparsely populate the space
2. Use Gaussian Process regression to model the χ^2 landscape
3. Populate regions that are likely to be useful

For our triaxial searches, we've used this customized routine and only needed 3000 – 5000 $\sim 4^6$ models across 3 iterations for 6 parameters. ($\sim 80,000$ CPU-hours)

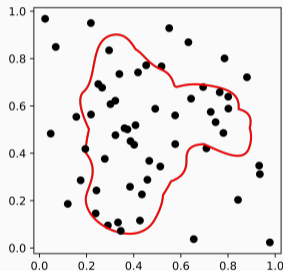


Our Strategy (inspired by Bayesian Optimization and nested sampling):

1. Sparsely populate the space
2. Use Gaussian Process regression to model the χ^2 landscape
3. Populate regions that are likely to be useful

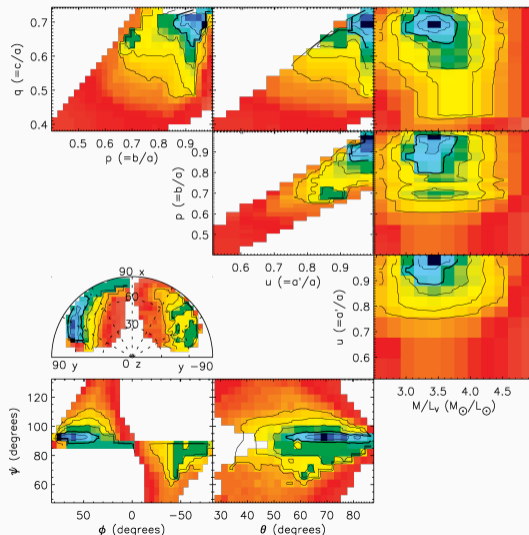
For our triaxial searches, we've used this customized routine and only needed 3000 – 5000 $\sim 4^6$ models across 3 iterations for 6 parameters. ($\sim 80,000$ CPU-hours)

For a reasonable-resolution grid search (10 pt per dimension), we'd need $O(10^6)$ models – 20,000,000 CPU-hours!

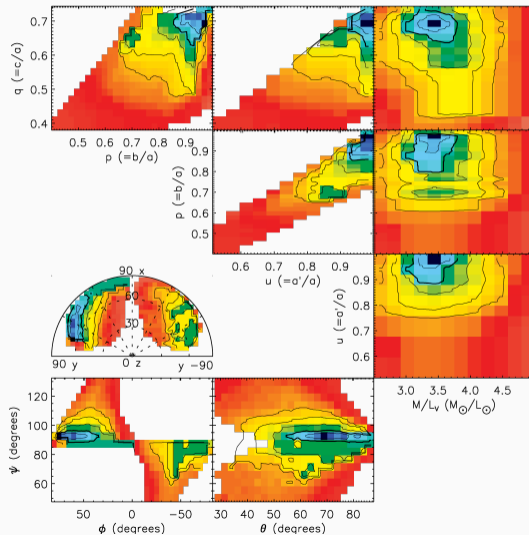


- The 3D shape is determined through **deprojection** of the 2D surface brightness profile (we use MGEs)
- This deprojection requires the choice of 3 parameters – viewing angles (θ, ϕ, ψ) or axis ratios (u, p, q) .

- The 3D shape is determined through **deprojection** of the 2D surface brightness profile (we use MGEs)
- This deprojection requires the choice of 3 parameters – viewing angles (θ, ϕ, ψ) or axis ratios (u, p, q).
- Not all choices of these parameters produce valid deprojections ($0 \leq q \leq uq' \leq p \leq u \leq 1$)



- The 3D shape is determined through **deprojection** of the 2D surface brightness profile (we use MGEs)
- This deprojection requires the choice of 3 parameters – viewing angles (θ, ϕ, ψ) or axis ratios (u, p, q) .
- Not all choices of these parameters produce valid deprojections $(0 \leq q \leq uq' \leq p \leq u \leq 1)$
- We've found an additional set of parameters which map the deprojectible shape space to a unit cube with minimal covariances



$$T = \frac{1-p^2}{1-q^2} \quad T_{\text{maj}} = \frac{1-u^2}{1-p^2} \quad T_{\text{min}} = \frac{(uq')^2 - q^2}{p^2 - q^2}$$

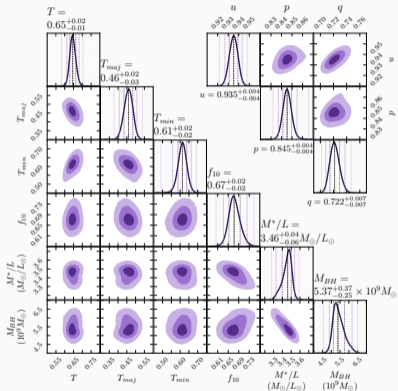
Our Observations

Our Data

Triaxiality!

Triaxial Schwarzschild Modelling

Results!



M87 Property (units)

Inferred value

Black hole mass M_{BH} ($10^9 M_{\odot}$)

$5.37^{+0.37}_{-0.25} \pm 0.22$

Inner M^*/L (V-band; M_{\odot}/L_{\odot})

$8.65^{+0.10}_{-0.15} \pm 0.38$

Dark matter fraction at 10 kpc f_{10}

0.67 ± 0.02

Shape parameter T

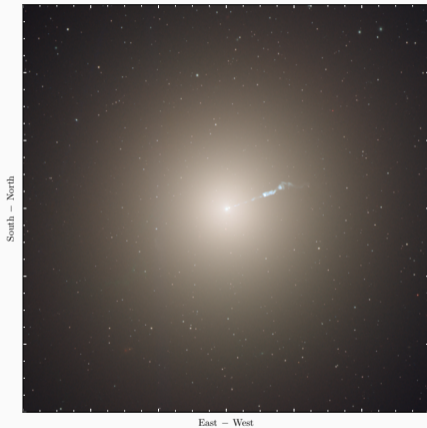
0.65 ± 0.02

Average middle-to-long axis ratio p

0.845 ± 0.004

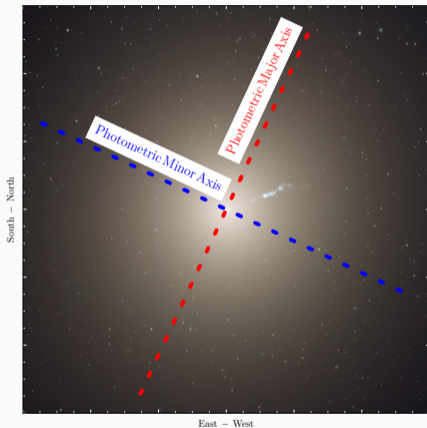
Average short-to-long axis ratio q

0.722 ± 0.007

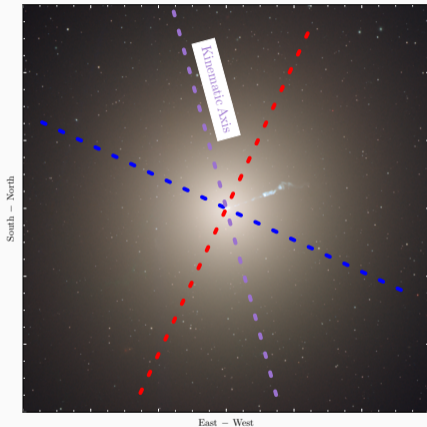


PA on Sky
($^{\circ}$ E of N)

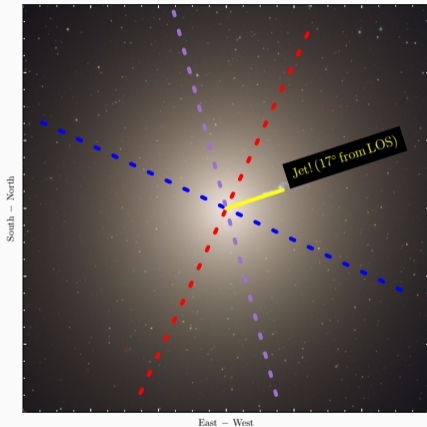
Angle from
Line of Sight



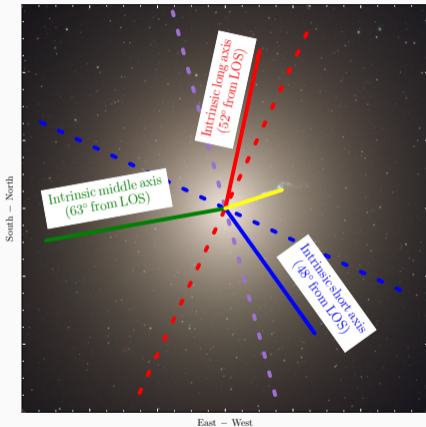
	PA on Sky (° E of N)	Angle from Line of Sight
Photometric Major Axis	-25°	—
Photometric Minor Axis	+65°	—



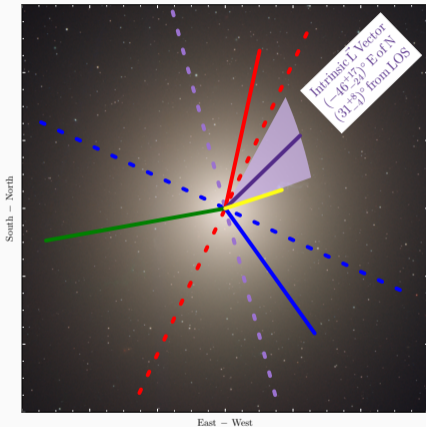
	PA on Sky (° E of N)	Angle from Line of Sight
Photometric Major Axis	-25°	—
Photometric Minor Axis	+65°	—
Kinematic Axis	-165°	—



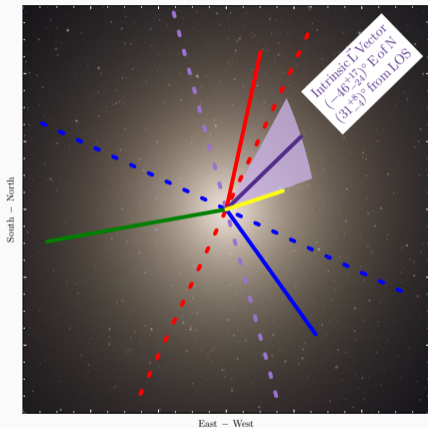
	PA on Sky (° E of N)	Angle from Line of Sight
Photometric Major Axis	-25°	—
Photometric Minor Axis	+65°	—
Kinematic Axis	-165°	—
Jet!	-72°	17°



	PA on Sky (° E of N)	Angle from Line of Sight
Photometric Major Axis	-25°	—
Photometric Minor Axis	+65°	—
Kinematic Axis	-165°	—
Jet!	-72°	17°
Intrinsic Long Axis	-12°	52°
Intrinsic Middle Axis	+100°	63°
Intrinsic Short Axis	-144°	48°



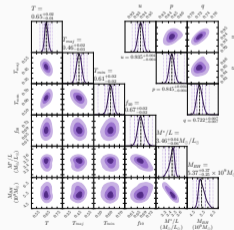
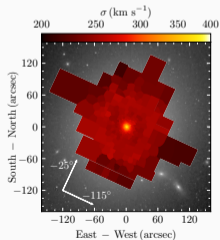
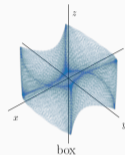
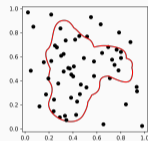
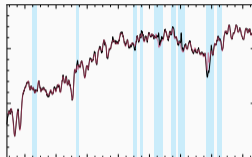
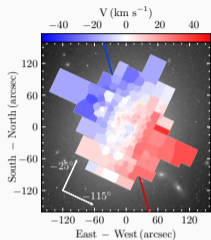
	PA on Sky (° E of N)	Angle from Line of Sight
Photometric Major Axis	-25°	—
Photometric Minor Axis	+65°	—
Kinematic Axis	-165°	—
Jet!	-72°	17°
Intrinsic Long Axis	-12°	52°
Intrinsic Middle Axis	+100°	63°
Intrinsic Short Axis	-144°	48°
Intrinsic \vec{L} Vector (between 80" and 150")	$(-46^{+17}_{-24})^\circ$	$(31^{+7}_{-4})^\circ$



	PA on Sky (° E of N)	Angle from Line of Sight
Photometric Major Axis	-25°	—
Photometric Minor Axis	$+65^\circ$	—
Kinematic Axis	-165°	—
Jet!	-72°	17°
Intrinsic Long Axis	-12°	52°
Intrinsic Middle Axis	$+100^\circ$	63°
Intrinsic Short Axis	-144°	48°
Intrinsic \vec{L} Vector (between 80" and 150")	$(-46^{+17}_{-24})^\circ$	$(31^{+7}_{-4})^\circ$

The intrinsic angular momentum axis of M87's stellar component is only $(17^{+11}_{-7})^\circ$ from the jet!

Thank you! (Questions?)



$M_{\text{BH}} (10^9 M_{\odot})$

Shape parameter T

Axis ratio p

Axis ratio q

$5.37^{+0.37}_{-0.25} \pm 0.22$

0.65 ± 0.02

0.845 ± 0.004

0.722 ± 0.007