

# A Duet of Black Holes from the TriOS (Triaxial Orbit Superposition) Code

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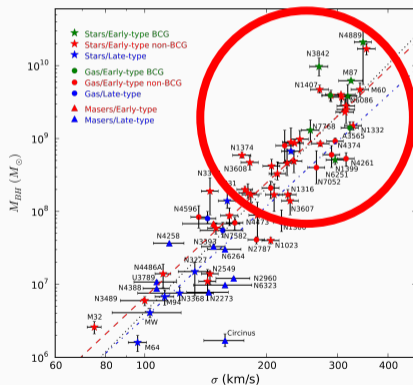
Emily Liepold  
emilyliepold@berkeley.edu

UC Berkeley

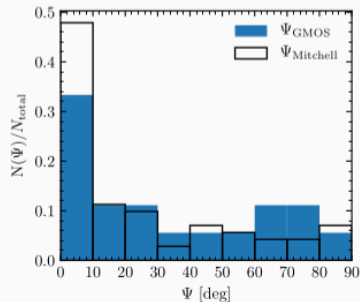


# Motivation: What are we looking at?

The **MASSIVE** Survey targets **MASSIVE** galaxies with **MASSIVE** black holes



- These galaxies often have kinematic misalignments
- Kinematic misalignments strongly suggest a **triaxial** intrinsic shape (not axisymmetry!)



## Motivation: Why do we care about the shape?

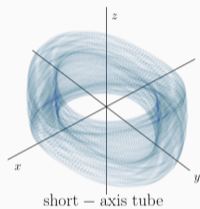
Shape of  $\rho \rightarrow$  Shape of  $\Phi \rightarrow$  Symmetries of  $\Phi \rightarrow$  Conserved quantities and allowed orbits

Symmetry		Conserved Quantity	Orbits
Spherical	$\frac{d\Phi}{d\Omega} = 0$	$(E, \vec{L})$	Rosettes in fixed planes
Axisymmetry	$\frac{d\Phi}{d\phi} = 0$	$(E, L_z, I_3)$	Loops about symmetry axis
Triaxiality	Eh...	$(E, I_2, I_3)$	It's complicated...

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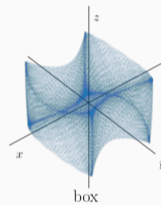
## Loop Orbits

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## Box Orbits

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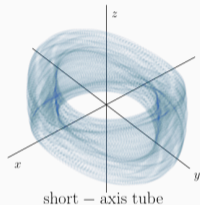


# Orbits in triaxial potentials

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## Loop Orbits

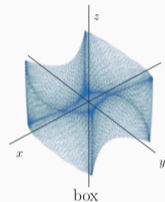
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Appears in axisymmetric potentials

## Box Orbits

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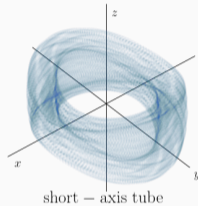


Not present in axisymmetry!

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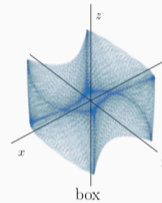
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Appears in axisymmetric potentials  
Persistent sense of rotation about  
either the **short** or **long** axis

## Box Orbits

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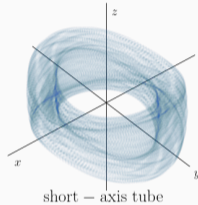
Not present in axisymmetry!  
No persistent sense of rotation

# Orbits in triaxial potentials

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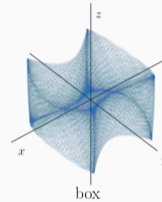
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Appears in axisymmetric potentials  
Persistent sense of rotation about  
either the **short** or **long** axis  
**Centrophobic**

## Box Orbits

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Not present in axisymmetry!  
No persistent sense of rotation  
Can be **Centrophilic**



Schwarzschild 1979: Can triaxial stellar systems in dynamical equilibrium be self-consistent?

Strategy:

1. Propose a (triaxial) stellar density distribution
2. Integrate representative orbits that span the phase space
3. Superimpose those orbits such that (1) is reproduced

# The TriOS Triaxial Orbit Superposition Code

van den Bosch+ 2008: Development of a **fortran**-based code for Schwarzschild orbit modelling in triaxial stellar potentials.

Model includes BH, stars, and dark matter halo:

$$\Phi = \Phi_{BH} + \Phi_* + \Phi_{DM}$$

Stellar kinematics (LOSVDs) described by Gauss-Hermite expansion with  $y = (v - V)/\sigma$ :

$$f(v) = \frac{e^{-\frac{y^2}{2}}}{\sqrt{2\pi\sigma^2}} \left[ 1 + \sum_{m=3}^n h_m H_m(y) \right]$$

2D (projected) and 3D (intrinsic) mass distributions are constrained for self-consistency. The code was un-named. We call our improved version 'TriOS' (**T**riaxial **O**rbit **S**uperposition)

We've made a number of changes to the original **vdB+08** code

- We've added a mode which **axisymmetrizes** the orbits, effectively making TriOS an (optionally) axisymmetric code.
- Orbits in triaxial potentials were **improperly mirrored** in the original code. We've fixed this
- Orbits near the BH's sphere of influence precess **slowly**. Orbits must be integrated for up to  $\sim 2000$  dynamical times before we find model convergence (vs 200 in previous usage)
- The original **orbit sampling density** leads to spurious biases in the preferred triaxiality parameter  $T$ . Doubling the sampling density solves the issue in most cases.
- Re-writing and tuning the routines for **PSF convolution** and **acceleration pre-caching** sped up the code by 5 – 10 $\times$  overall.
- In some circumstances, **energy conservation** was not checked after orbit integration or **mass self-consistency** was ill-enforced. These have been fixed.

Each **TriOS** model gives a  $\chi^2$  value for a single point in the parameter-space

- We need to search over  $M_{BH}$ ,  $M/L$  (1 or 2 parameters), shape (3 parameters), and halo (1 or 2 parameters) – at least **6-8 dimensions**. (Grid Searches are inefficient)

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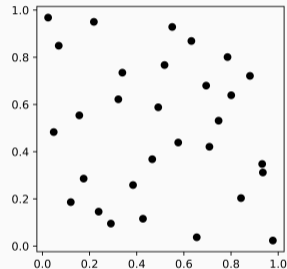
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- As data improves, confidence volumes **shrink** with  $\sim (\text{Number of Constraints})^{-D/2}$

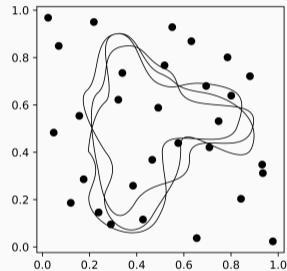
Our Strategy (inspired by Bayesian Optimization and nested sampling):

1. Sparsely populate the space



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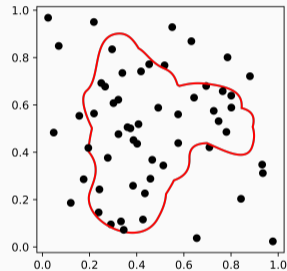
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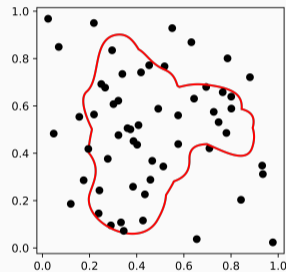
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For our triaxial searches, we used this customized routine and only needed 3000 – 5000  $\sim 4^6$  across 3 iterations for a 6D search. ( $\sim 80,000$  CPU-hours)



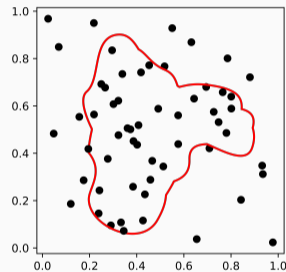
# Efficient Sampling for Triaxial Modelling (Liepold+ future paper)

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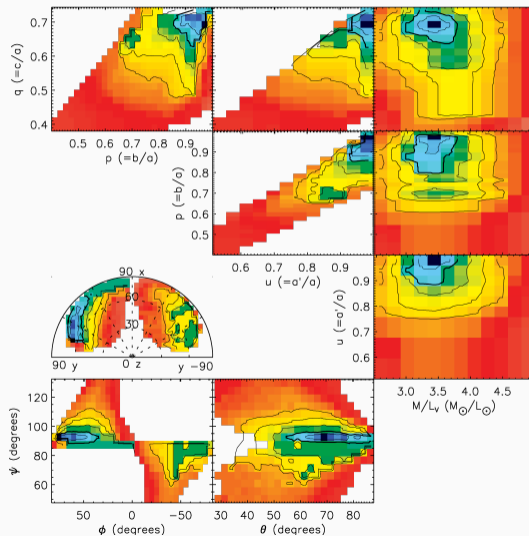
For a reasonable-resolution grid search (10 pt per dimension), we'd need  $O(10^6)$  models – 20,000,000 CPU-hours!



- The 3D shape is determined through **deprojection** of the 2D surface brightness profile (we use MGEs)
- This deprojection requires the choice of 3 parameters – viewing angles  $(\theta, \phi, \psi)$  or axis ratios  $(u, p, q)$ .

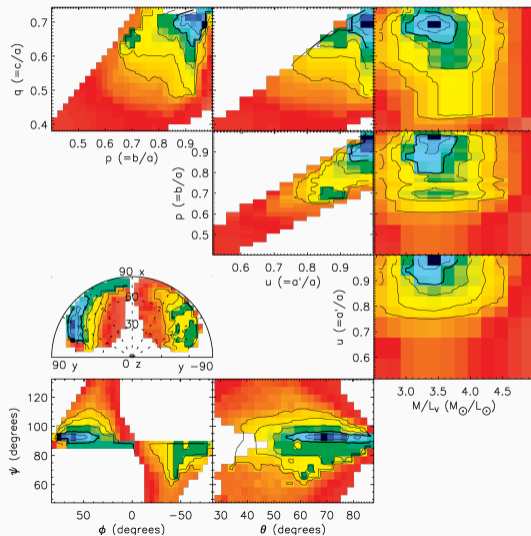
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- Not all choices of these parameters produce valid deprojections ( $0 \leq q \leq uq' \leq p \leq u \leq 1$ )



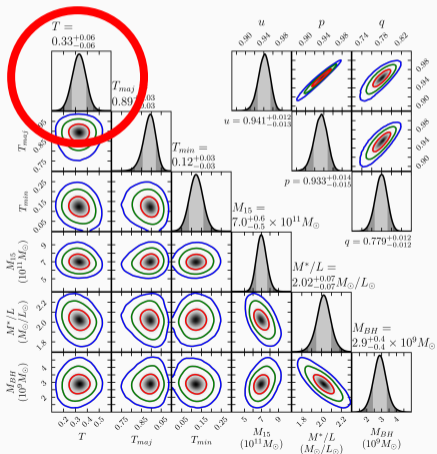
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- Not all choices of these parameters produce valid deprojections  $(0 \leq q \leq uq' \leq p \leq u \leq 1)$
- We've found an additional set of parameters which map the deprojectible shape space to a unit cube with minimal covariances

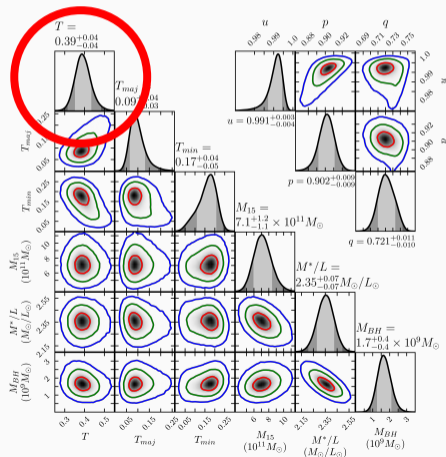


$$T = \frac{1-p^2}{1-q^2} \quad T_{\text{maj}} = \frac{1-u^2}{1-p^2} \quad T_{\text{min}} = \frac{(uq')^2 - q^2}{p^2 - q^2}$$

## NGC1453

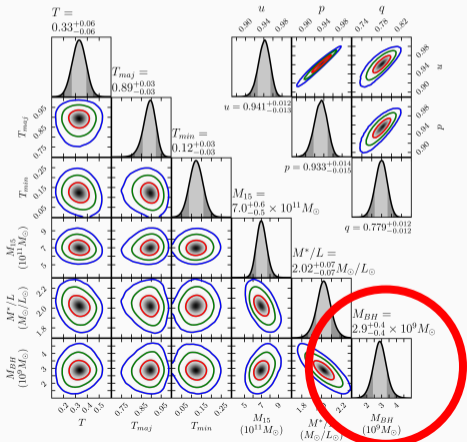


## NGC2693

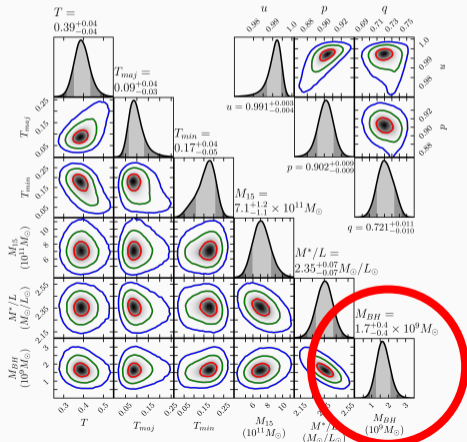


# Triaxial NGC1453 and NGC2693 (Quenneville, Liepold, and Ma 2021b), (Pilawa, Liepold+22)

## NGC1453

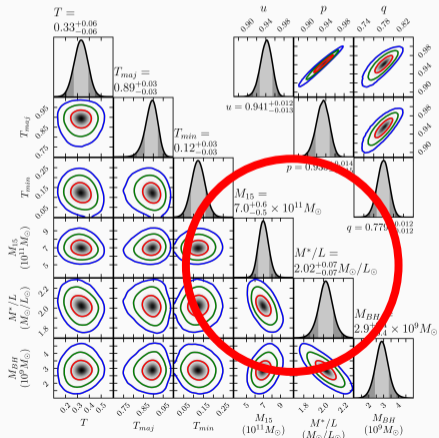


## NGC2693

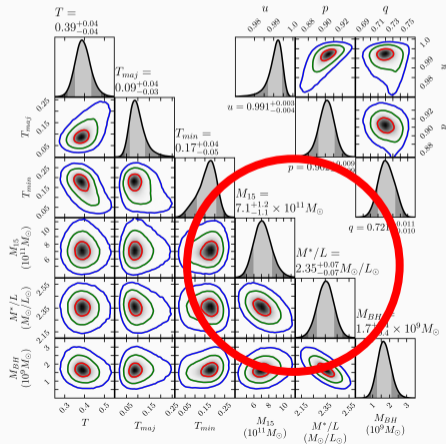




## NGC1453

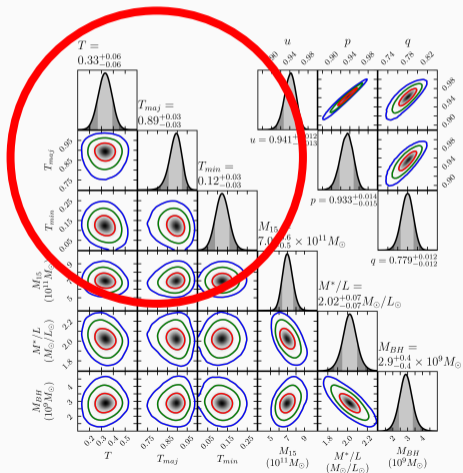


## NGC2693

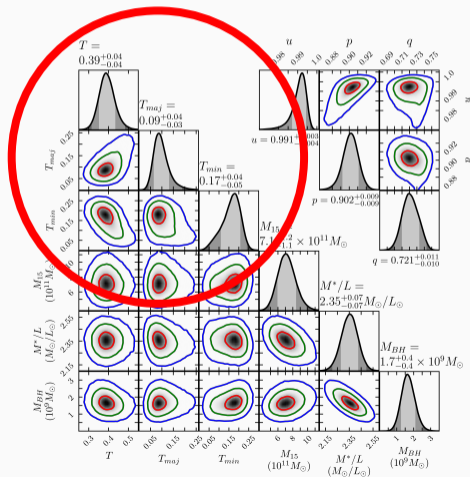


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## NGC1453



## NGC2693



## Looking Forward: Model Complexity

- Recent papers (Lipka+Thomas 2021+2022) suggest that the **model complexity** must be taken into account while finding the best-fit models by adding a penalty term to the model  $\chi^2$
- Our preliminary tests suggest that this issue is **significant for axisymmetric models** and a bias in inclination is present if the complexity is ignored
- Our preliminary tests find that the issue is **far less important for triaxial models** and the preferred shape is more-or-less unchanged when reasonable penalty terms are introduced.

