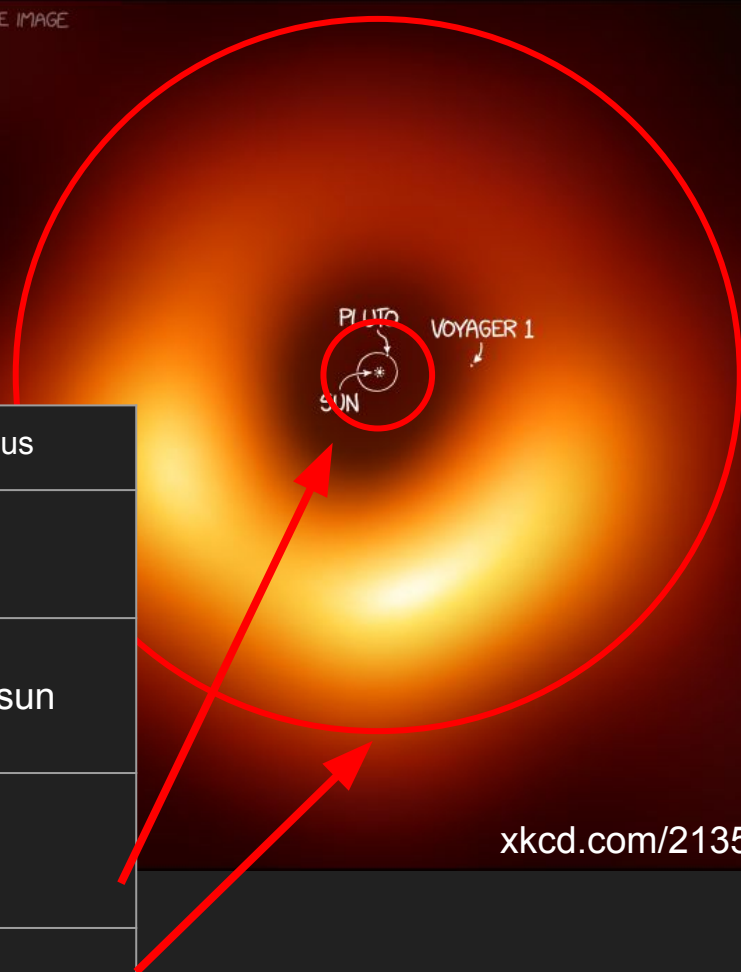


A Search for the Largest Supermassive Black Holes

Supermassive Black Holes

- Unimaginably huge
- Live in centers of most large galaxies

	Mass	Event Horizon Radius
Sun	$1 M_{\text{sun}}$	Walk around campus
Milky Way Sgr A*	4 Million M_{sun}	3 x Circumference of sun
M87	6 Billion M_{sun}	3 x distance to Pluto
Holmberg 15A	40 Billion M_{sun}	Huge!



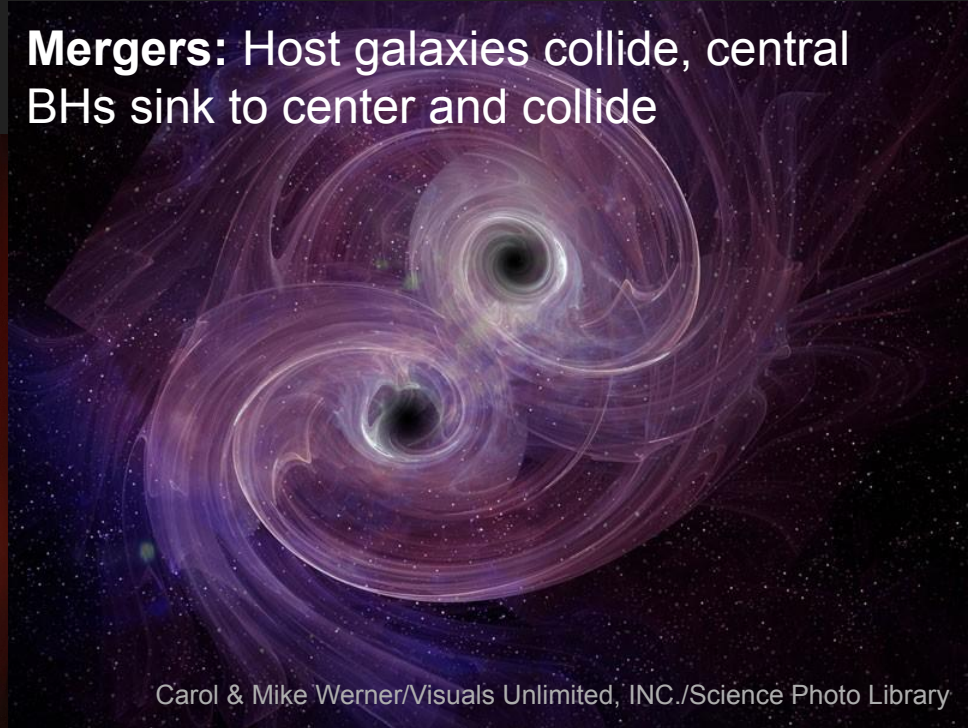
Evolution of SMBHs

- Supermassive black holes become supermassive through **accretion** and **mergers**

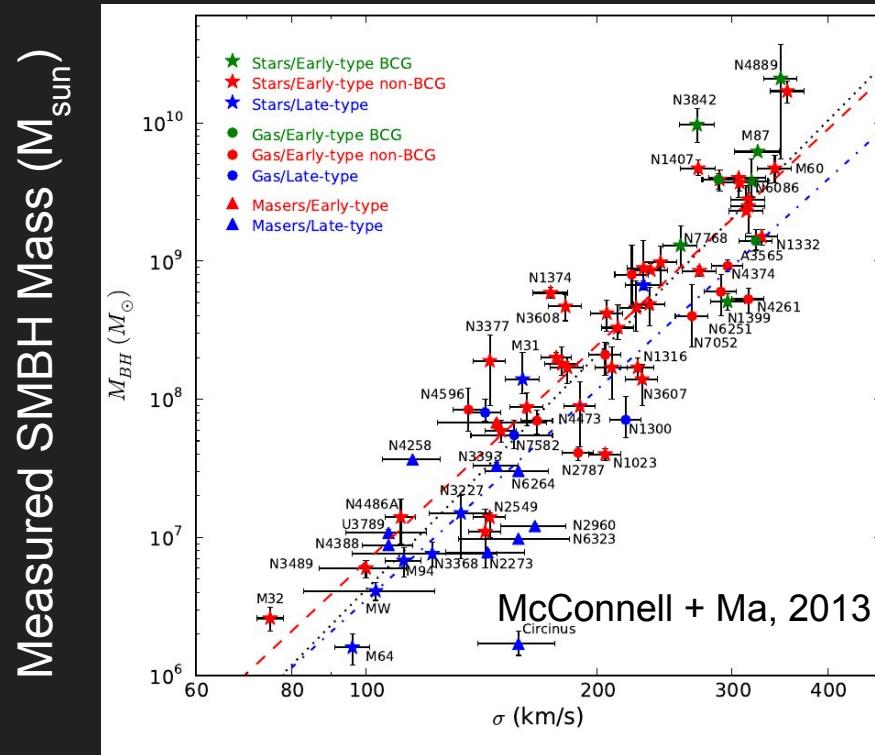
Accretion: Gas falls onto the central BH



Mergers: Host galaxies collide, central BHs sink to center and collide

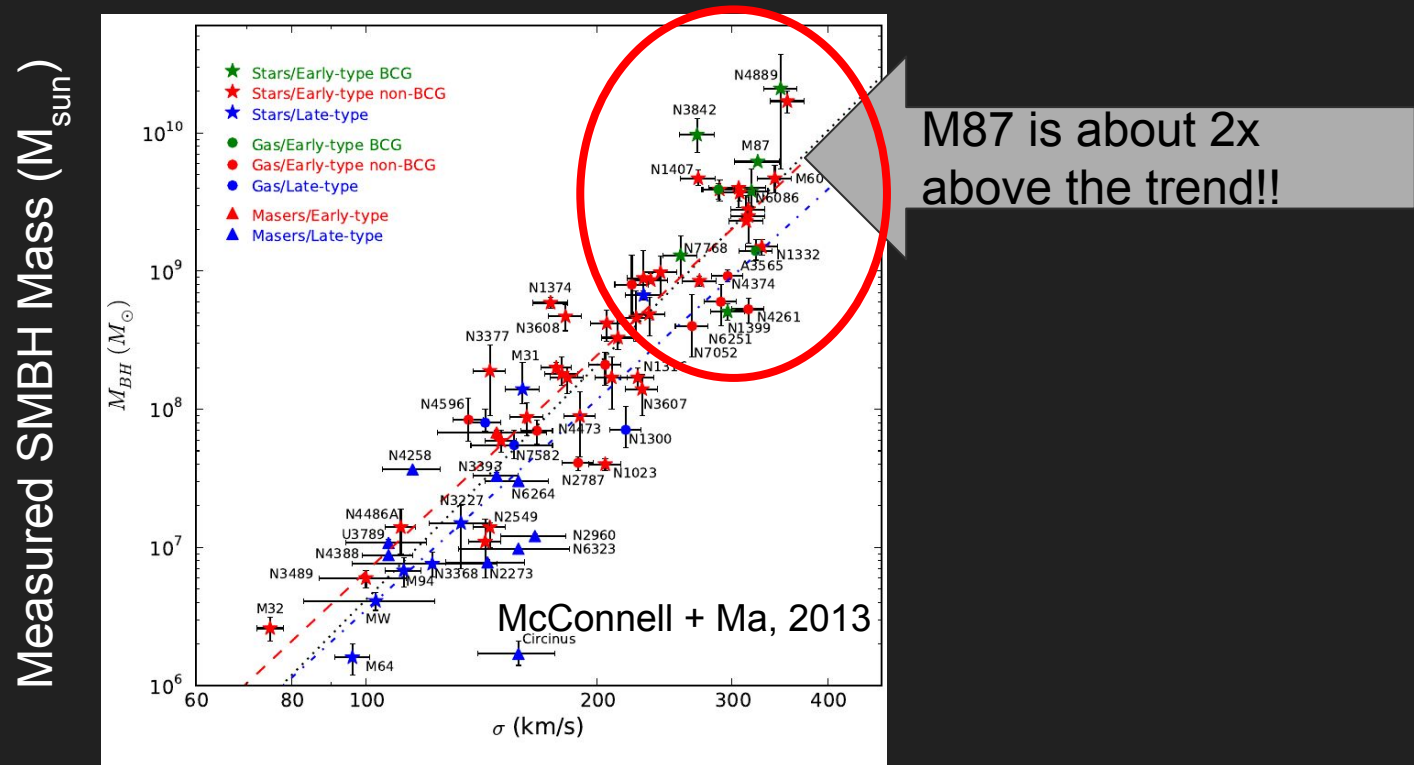


SMBHs co-evolve with their host galaxies



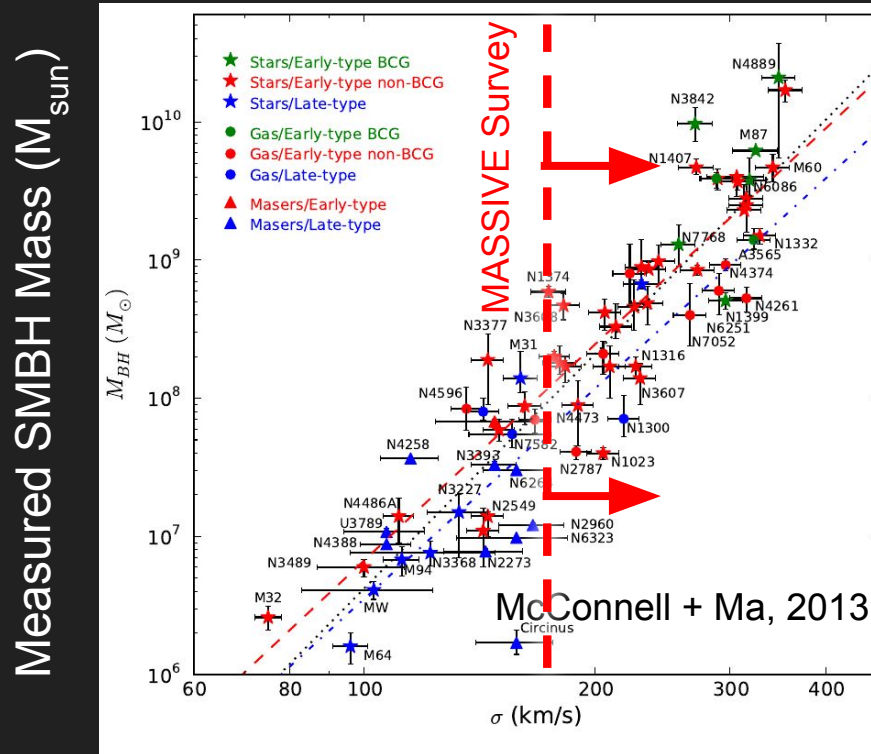
σ ~ speed of stars far away from BH ~ tracer of galaxy mass

SMBHs co-evolve with their host galaxies



$\sigma \sim$ speed of stars far away from BH \sim tracer of galaxy mass

SMBHs co-evolve with their host galaxies



MASSIVE Survey explores ~ 100 most massive early type galaxies in our neighborhood

(Note: MASSIVE selects targets by M_* , not M_{BH} , so vertical line is not representative of real cuts)

$\sigma \sim$ speed of stars far away from BH \sim tracer of galaxy mass

Finding Black Holes: Observables (stars!)



Photometry (photos)

Where are the stars?

Finding Black Holes: Observables (stars!)

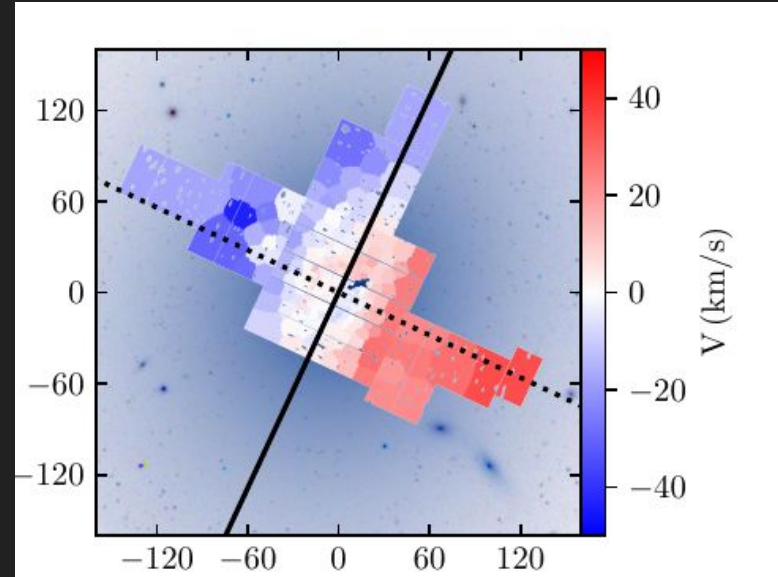


Photometry (photos)

Where are the stars?

Kinematics (spectra)

How do the stars move?

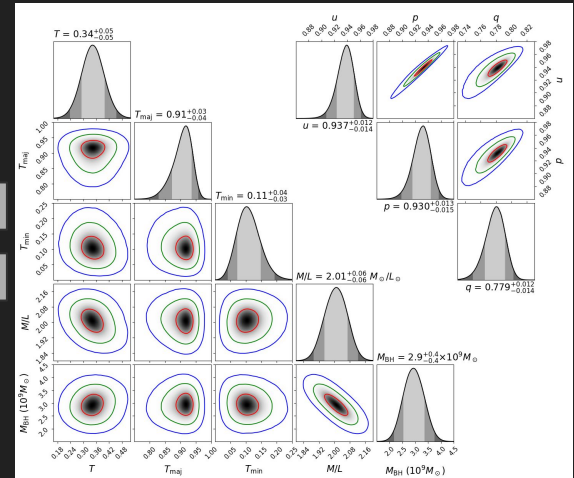
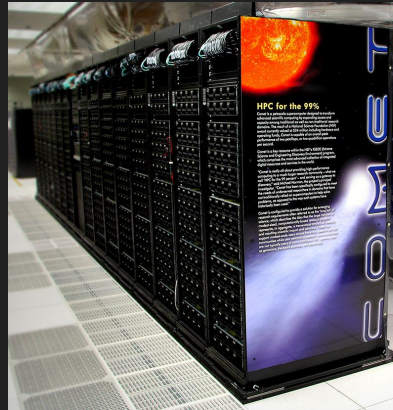


Liepold+ (in prep)

- We only measure quantities in **projection** -- 3D positions and velocities can't uniquely be reconstructed from the data

Finding Black Holes: Modelling

1. Assume a **deprojection** of the photometry and a **mass model**
2. Integrate **>100k** representative stellar orbits in that potential
3. Combine those orbits to match the measured kinematics
4. Repeat 1-3 to find the best mass model and deprojection



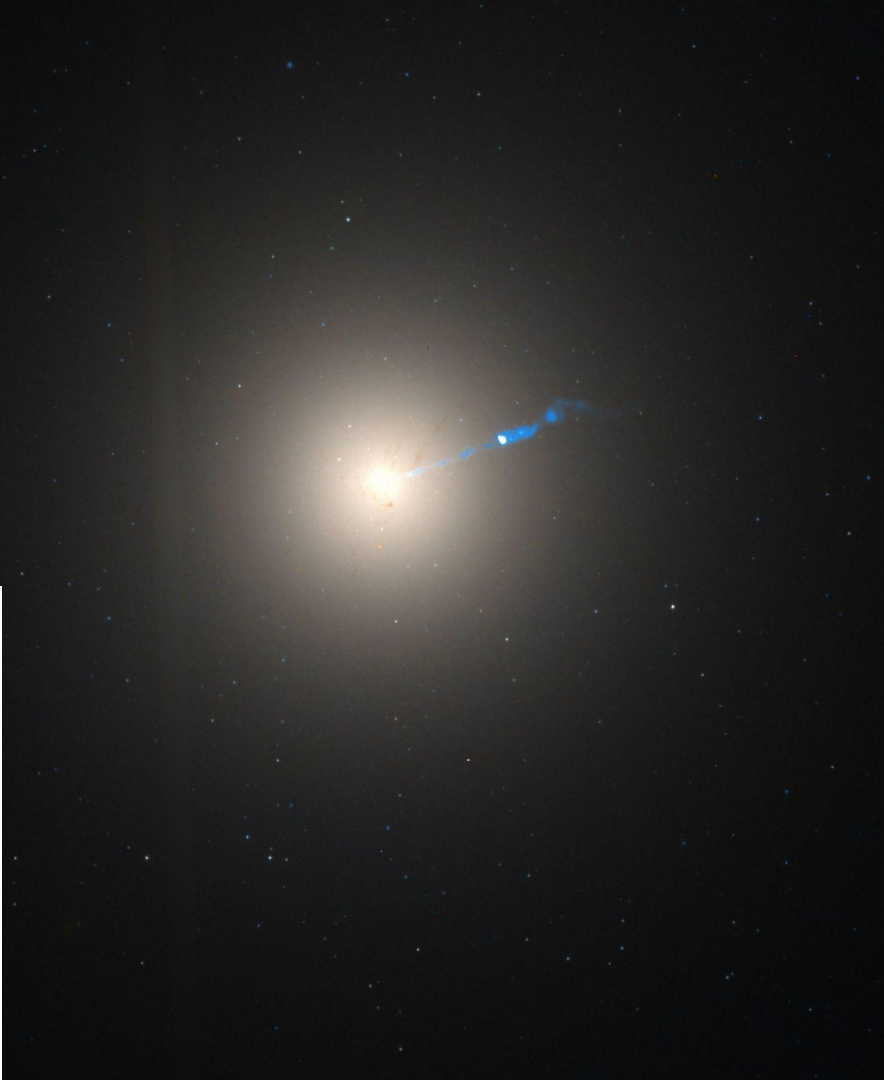
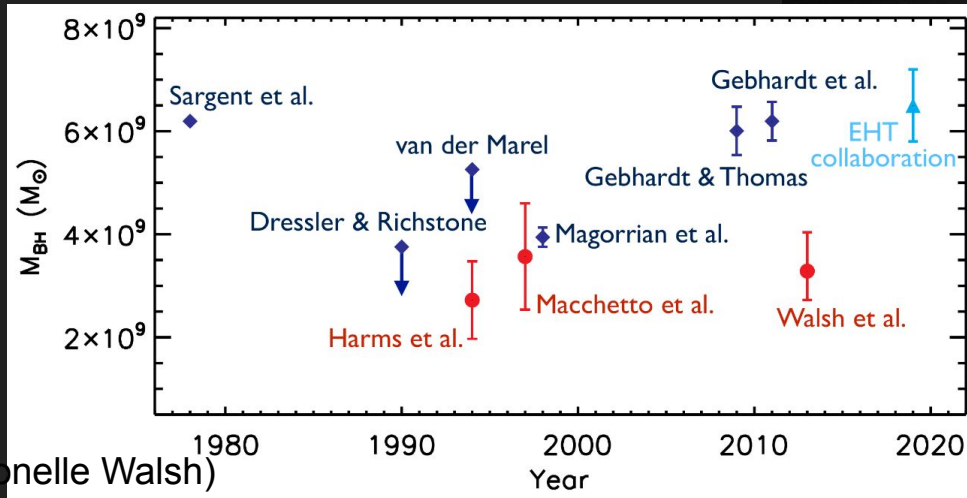
Digression: Deprojecting Ellipses

- Elliptical galaxies are Elliptical!
- 3D ellipsoids look like ellipses when projected on the sky
- In general, 3D ellipsoids can be **triaxial** with 3 different axis lengths
- Mis-modelling the 3D shape can **significantly** bias the measured BH mass



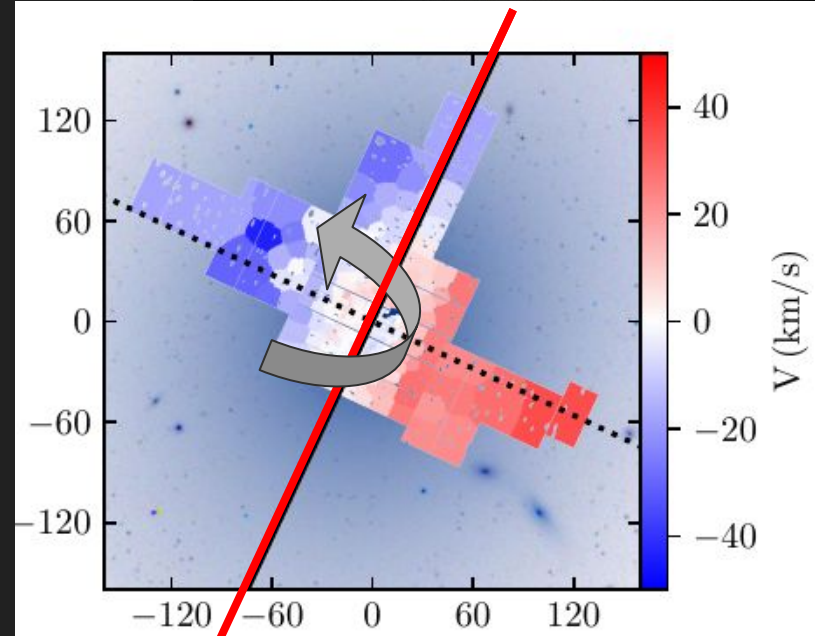
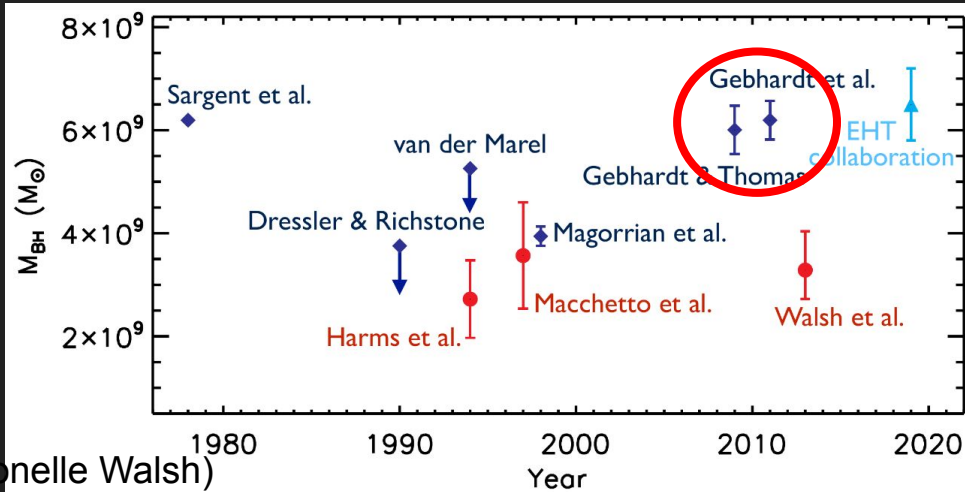
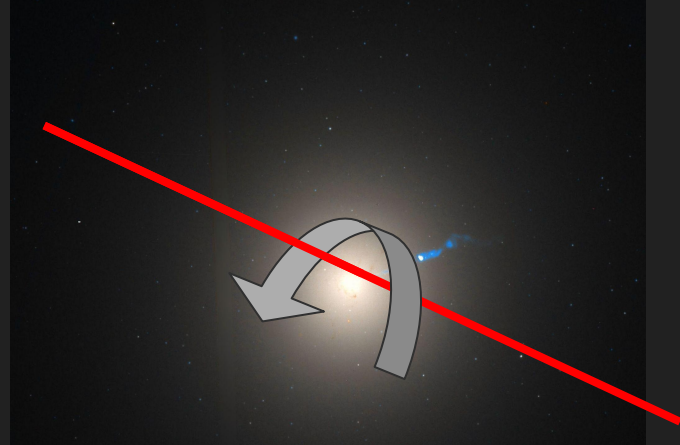
M87: A (brief) history

- (Axisymmetric) Orbit modelling suggests $M_{\text{BH}} \sim 6 \times 10^9 M_{\text{sun}}$
- Gas modelling suggests $M_{\text{BH}} \sim 3 \times 10^9 M_{\text{sun}}$
- EHT suggests $M_{\text{BH}} \sim 6 \times 10^9 M_{\text{sun}}$



M87's puzzling shape

- Axisymmetric modelling **only** allows rotation about the symmetry axis
- Our recent observations find that the rotation is 90 degrees misaligned!
- M87 is not axisymmetric!
- Is the mass really **6 Billion**?



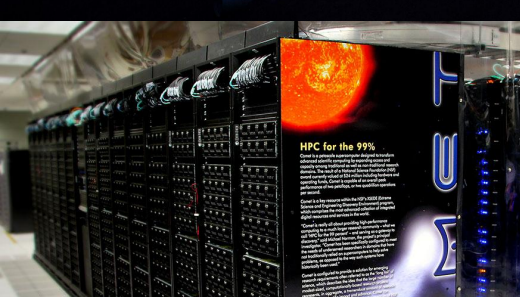
The Instruments



Keck



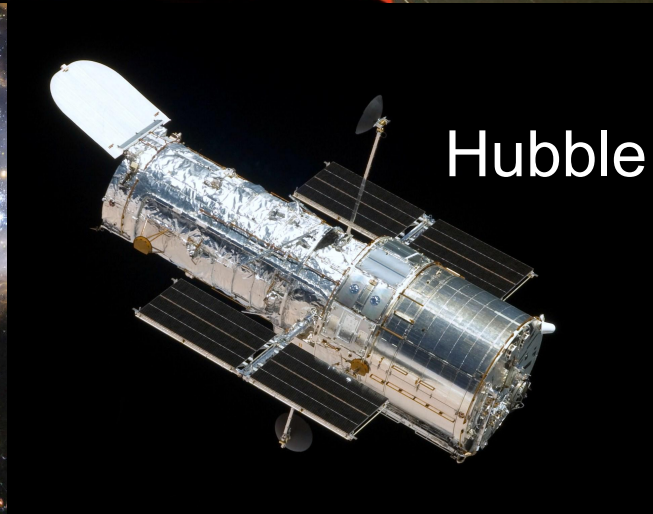
Gemini



San Diego
Supercomputer Center
(UCSD)

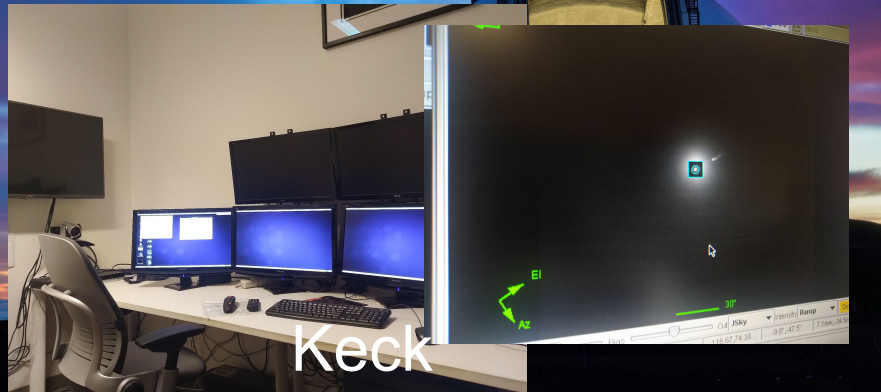


James Webb



Hubble

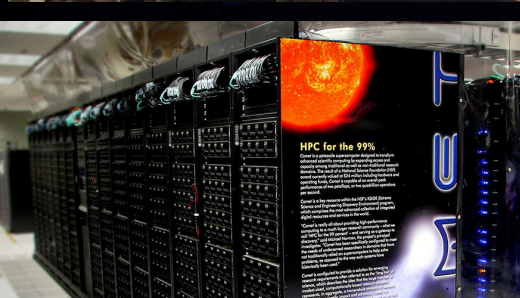
The Instruments



Keck



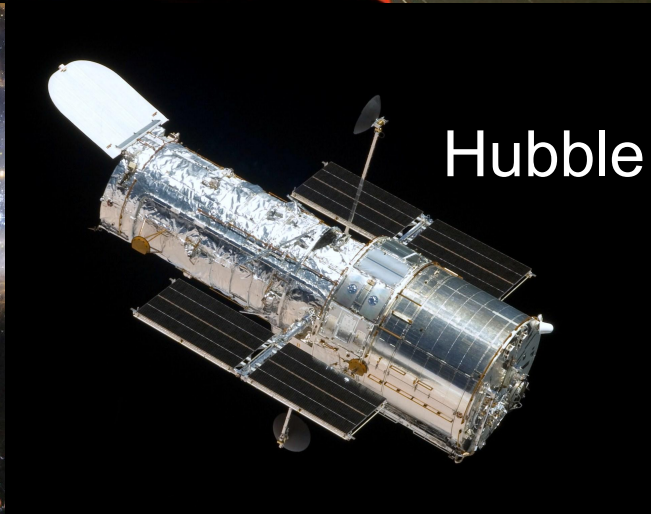
Gemini



San Diego
Supercomputer Center
(UCSD)

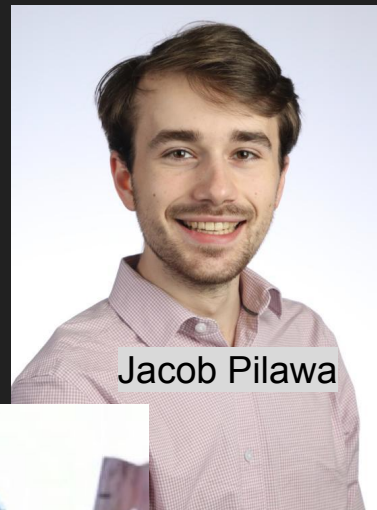


James Webb



Hubble

The People



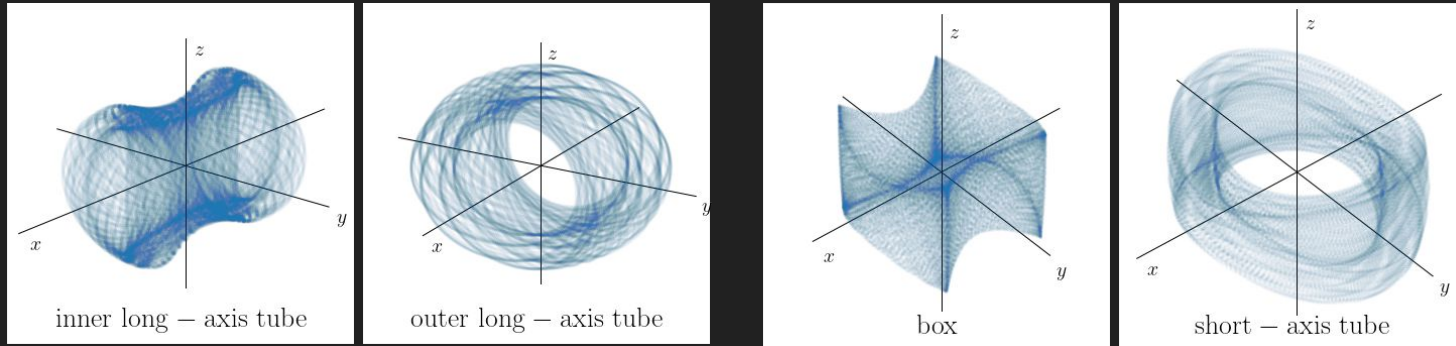
Myself!

Questions?

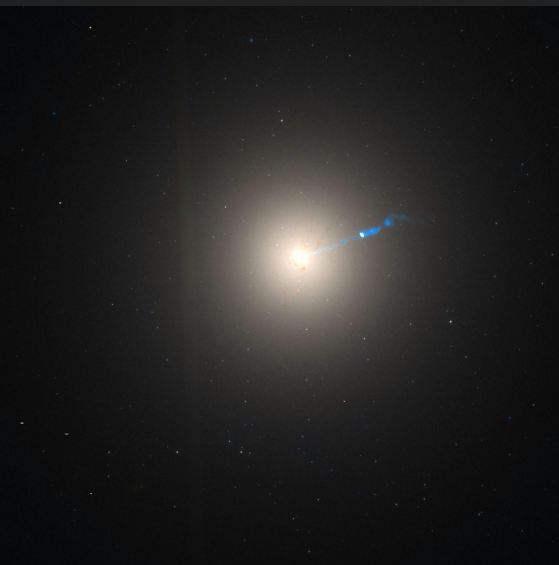
Additional slides (TBD)

The importance of shapes

- **Axisymmetric** models only allow orbits about the symmetry axis
- **Triaxial** models allow orbits around **short** or **long** axis and orbits which pass through center (**boxes**)
- **Box** orbits can mimic the kinematic signatures of a black hole
- Mis-modelling the 3D shape can bias the BH measurement by a factor of >2



Finding Black Holes: Observables (stars!)

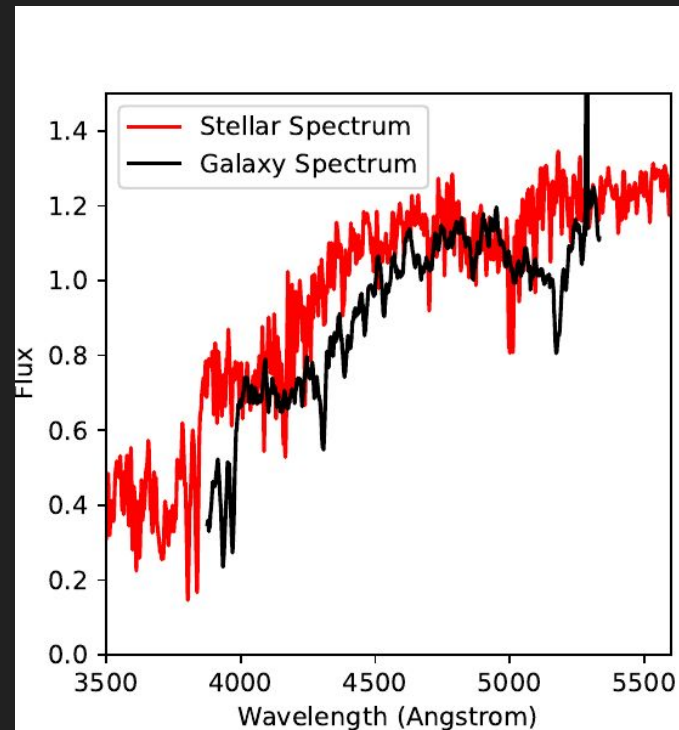


Photometry (photos)

Where are the stars?

Kinematics (spectra)

How do the stars move?



Finding Black Holes: Observables (stars!)

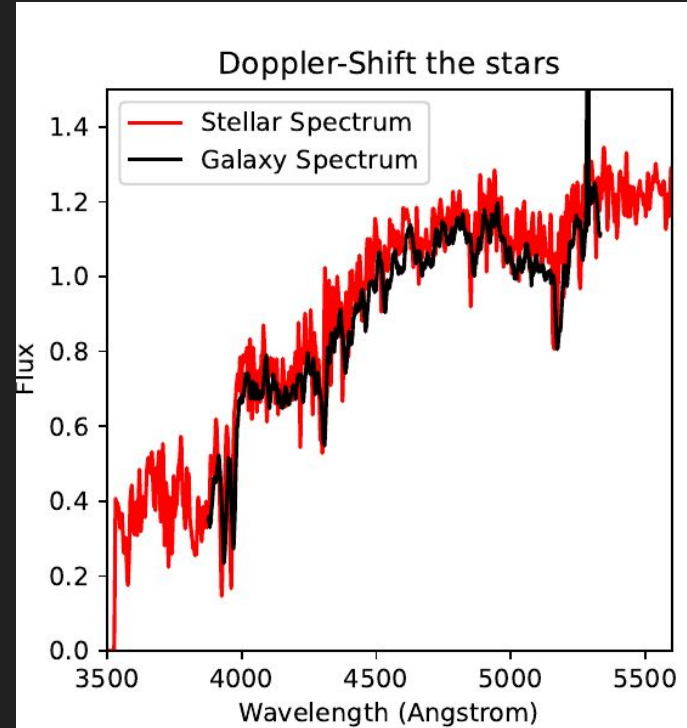


Photometry (photos)

Where are the stars?

Kinematics (spectra)

How do the stars move?



Finding Black Holes: Observables (stars!)



Photometry (photos)

Where are the stars?

Kinematics (spectra)

How do the stars move?

